COLUMBIA UNIVERSITY

IN THE CITY OF NEW YORK

Abstract

- 5G-and-beyond wireless networks will utilize millimeter-wave (mmWave) frequencies to achieve significantly higher data rates [1].
- The mmWave radio frequency signal experiences high path loss especially into buildings.
- Following from prior outdoors measurements [2, 3], to quantify and understand building penetration by a 28 GHz mmWave signal, a street-based measurement campaign in the deployment area of the PAWR COSMOS testbed in New York City was conducted. Results for two distinct types of common buildings are presented; newer buildings primarily constructed with glass and metal, and older buildings primarily constructed with brick.
- Path gain values with their fitted lines and the effective azimuthal beamforming gains are computed.
- Results can inform COSMOS testbed development, including the deployment of IBM 28 GHz phased array antenna modules [4] as well as indoor 5G/6G wireless access in locations such as public schools.

COSMOS Testbed

- Cloud Enhanced Open Software Defined Mobile Wireless Testbed for *City-Scale Deployment* (COSMOS) is a city-scale programmable testbed for advanced wireless technologies in West Harlem, New York City [5, 6].
- The COSMOS Team includes Rutgers, Columbia, NYU, University of Arizona and CCNY, alongside city, community and industry collaborations with New York City, Silicon Harlem, and IBM.





COSMOS' multi-layered computing architecture



Blue – Battery and RF circuits **Green** – Omni-directional antenna **Green** – 10 degree horn antenna



Red – Base and rotating platform Blue – Power meter and Raspberry Pi

- We use a 28 GHz portable narrowband channel sounder.
- The transmitter (Tx) is equipped with an omni-directional antenna, and the receiver (Rx) is equipped with a 10-degree horn antenna with 24.5 dBi gain mounted on a rotating platform.
- Rx records power measurements at a rate of 740 samples per second using an onboard Raspberry Pi that is wirelessly controlled by a PC.

Outdoor-to-Indoor 28 GHz mmWave Measurements in the COSMOS Testbed Deployment Area NOKIA Gulnur Avci^{1*}, Sienna Brent^{1*}, Sabbir Hossain^{2*}, Jared Moser^{3*}, Angel Daniel Estigarribia¹, Manav Kohli¹, Igor Kadota¹, Abhishek Adhikari¹, Dmitry Chizhik^{4*}, Jinfeng Du^{4*}, Rodolfo Feick⁵, Reinaldo Valenzuela⁴, Gil Zussman¹ ¹Columbia University, ²City College of New York, ³Stuyvesant High School, ⁴Nokia Bell Labs, ⁵Universidad Técnica Federico Santa María; * indicates equal contribution



Measurement Environment and Locations New-1, Rx height 25ft O New-2, Rx height 35ft Old-1, Rx height 14ft Old-2, Rx height 0-30ft Tx path on street COSMOS Testbed Area



Case Study – Hamilton Grange Middle School at Old-1

- Noticeable difference when moving Rx along red paths at Parking Lot (PL) and Basketball Courts (BC).
- 10dB difference in path loss at ~85m distance demonstrates the impact of the angle of incidence.
- ~85m distance at BC is at a low angle of incidence, ~85m distance at PL is at high angle of incidence.
- Implications on base station deployments: may need to avoid high angles of incidence.





Over 1,100 links measured across four different measurement sites, emulating different indoor environments and building types:

- **New-1** Corner office in modern building
- **New-2** Coffee shop in modern building **Old-1** – Classrooms in older public school
- **Old-2** Library and cafeteria in older university

• New-1 and New-2 are modern buildings with predominantly glass and metal construction.

• Old-1 and Old-2 are older buildings with predominantly brick and concrete construction. Each measurement location except **Old-1** is in an urban street canyon. Measurements taken at Old-1 are performed in the grounds of a middle school where the surrounding buildings are further away. • These four deployment scenarios are common throughout NYC and other northeast US cities.

Angle of Arrival (AoA) for Classrooms at Old-1

- received from.
- provide maximum coverage indoors.
- line direction from Tx to Rx.
- amount of scattering.
- and beamsteering.
- N_{Rx} = -90 dBm. With worst-case would be in excess of 1.3 Gbps.

Ongoing & Future Work

- More extensive measurements in the ones in the COSMOS testbed area.
- Deployment of IBM phased arrays into COSMOS medium nodes [4].
- Utilize measurement data to analyze (FWA) deployment [7].



Medium Node

References

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• Angular spectra can also be analyzed to understand AoAs; the directions power is

• Understanding AoA can help optimize location of outdoor mmWave nodes to

• Whether transmitter is placed in front of each classroom, or a single transmitter is placed in the middle of the classrooms, the peak power received is in the straight-

• Angular spectra in general show a large

• This will reduce the beamforming gain achievable in such links, but also lessens the need for ultra-precise beamforming

• Define typical values for a mmWave link: Tx power P_{Tx} = 28 dBm, Tx gain G_{Tx} = 23 dBi, Rx gain G_{Rx} = 9 dBi, Rx noise floor

measured path gain of -130 dB, a signalto-noise ratio (SNR) of 20 dB is achieved. At 200 MHz bandwidth, indoor data rates



Angular spectra for transmitter placed in front of each classroom



Angular spectra for single transmitter location in middle of classrooms

 $SNR = P_{Tx} + G_{Tx} + PG + G_{Rx} - N_{Rx}$

current locations, as well as new performance of IBM phased arrays when used in a fixed-wireless access



Future measurement locations





IBM Phased Array Electronics

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